

Boundary state black holes

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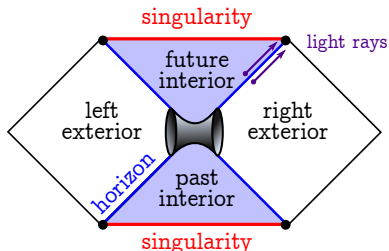
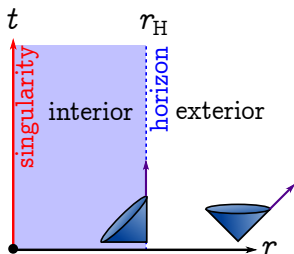


Based on 1810.10601 w/ Mark Van Raamsdonk, Moshe Rozali,
Sean Cooper, Chris Waddell (UBC), and Brian Swingle (UMD)

I. Black holes

Classical black holes

- Black hole: **region you can enter but can't leave.**[†]
- Light cone of infalling observer **tips over at horizon.**

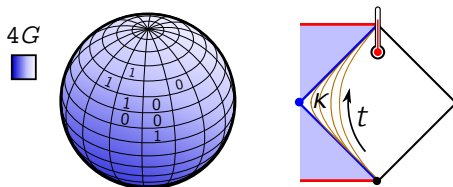


- Penrose diagram captures global structure.[‡] For spherically symmetric BH, get **two exteriors joined by wormhole.**

[†]Schwarzschild 1916. [‡]Penrose 1964; Carter 1966.

Black hole thermodynamics

- Remarkable fact: **black holes are thermodynamic systems**.[†]
- The **entropy is proportional to the horizon area**, $S = A/4G$.[‡]

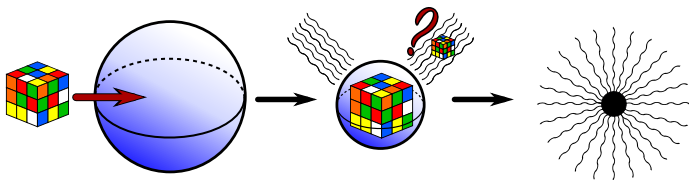


- **BHs emit Hawking radiation** at $T = \kappa/2\pi$, where κ is the surface gravity (proper acceleration at horizon).

[†]Bardeen, Carter and Hawking 1973. [‡]Hawking 1971; Bekenstein 1972. [§]Hawking 1974.

The information problem

- Radiation leads to a paradox: **BHs evaporate into thermal noise**. They destroy information about what fell in![†]

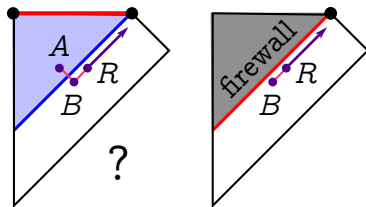


- Process is irreversible, like the second law. But **second law comes from ignoring microscopic details**.
- Suggests that **information is encoded in microscopic details of radiation**.[‡] Acts like a xerox machine on things falling in!

[†]Hawking 1975. [‡]Susskind, Thorlacius and Uglum 1993; Susskind and Thorlacius 1993.

Firewalls

- If radiation copies things falling in, **we get a paradox**. First, **copying entangles the black hole with radiation**.[†]
- Second, having smooth fields across the horizon requires **the horizon and interior to be entangled**.[‡]

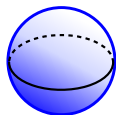


- Paradox: **only one can hold!**[§] “Monogamy” of entanglement. If horizon not smooth, **replaced by a high-energy “firewall”**.[‡]

[†]Page 1993. [‡]Unruh and Wald 1984. [§]Coffman, Kundu, and Wootters 1999; Mathur 2009. [‡]Almheiri, Marolf, Polchinski and Sully 2012.

Simulating the interior

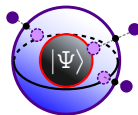
- Second option: “simulate” interior with stuff outside black hole.[†] Avoids monogamy issue.
- Use state $|\Psi\rangle$ outside BH as computational resource, giving state-dependent[‡] simulation.



classical



firewall



simulation

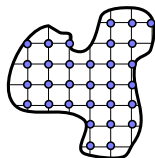
- Can't simulate everything behind horizon, so we expect a state-dependent amount of interior.[§]
- We will give precise realisation in AdS/CFT!

[†]Papadodimas and Raju 2014; Maldacena and Susskind 2013. [‡]Papadodimas and Raju 2015. [§]Shenker and Stanford 2013; de Boer, van Breukelen, Lokhande, Papadodimas and Verlinde 2018.

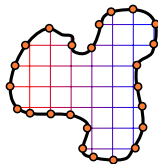
II. AdS/CFT

Gravity is holographic

- AdS/CFT is a theory of quantum gravity.
- Motivation: **quantum gravity is holographic**.[†] Unlike local theory, entropy in gravity scales with area rather than volume.



local QFT



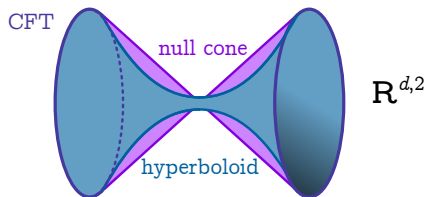
quantum gravity

- Simple argument: **BHs maximise entropy density**. Otherwise, you can collapse a system into a BH and violate second law.

[†]t Hooft 1993; Susskind 1995.

Matching symmetries

- AdS/CFT realises holography, with (AdS) gravity in $d + 1$ dimensions equal to (CFT) quantum theory in d dimensions.[†]
- AdS = anti-de Sitter space (constant negative curvature).
CFT = conformal field theory (conformally invariant QFT).

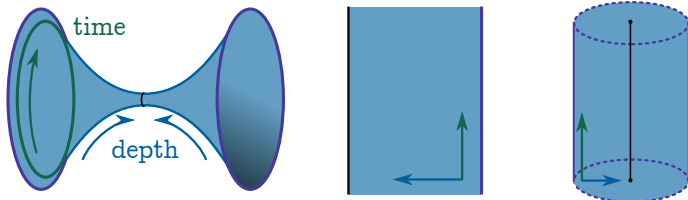


- We can embed AdS as hyperboloid $X^2 = L^2$ in $\mathbb{R}^{d,2}$. CFT lives on projective null cone $X^2 = 0$ in $\mathbb{R}^{d,2}$.[§]
- Symmetry group $SO(d, 2)$ on both sides matches!

[†]Maldacena 1997; Gubser, Klebanov and Polyakov 1998; Witten 1998. [§]Dirac 1935.

Pictures of AdS/CFT

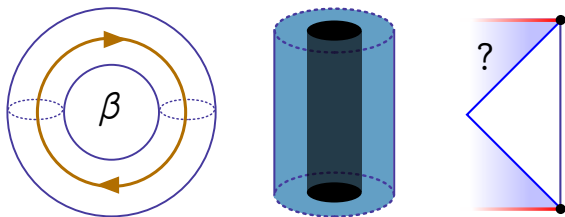
- Two important directions: **time on the hyperboloid is periodic**, so we unwrap it. **Depth is distance from purple boundary.**
- In time/depth coordinates, **Penrose diagram is a rectangle.**



- **CFT lives in flat space $\mathbb{R}^{d-1} \times \mathbb{R}$.** We can make space compact so that **CFT lives on a cylinder $\mathbb{S}^{d-1} \times \mathbb{R}$.**

Thermal states and black holes

- Empty AdS corresponds to CFT vacuum state. Now consider thermal state with temperature $T = 1/\beta$.
- System has period β in imaginary time,[†] so CFT cylinder is wrapped into a donut. In AdS, corresponds to a black hole![‡]

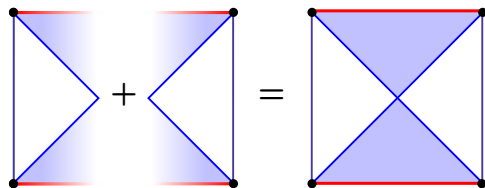


- Information problem: we don't know what's inside BH!

[†]Matsubara 1955. [‡]Hawking and Page 1983; Witten 1998.

Purification and wormholes

- Ignorance of BH interior corresponds to **fact that quantum state ρ is mixed**. What happens if we purify the state?
- Recipe for purifying ρ : **copy system, entangle copies, and apply $\sqrt{\rho}$** . Construction gives **thermofield double (TFD)**.

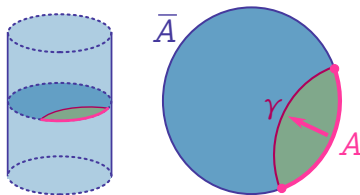


- Each copy has BH exterior. Natural to expect that **TFD is dual to a wormhole** with Schwarzschild AdS ends.[‡]

[†]Israel 1976. [‡]Maldacena 2001.

Entanglement and geometry

- **Geometry and entanglement connected.**[†] Ryu-Takayanagi (RT) formula gives similar connection.[‡]
- Pick a subsystem A of the CFT. “Push” A into bulk surface γ of minimal area.



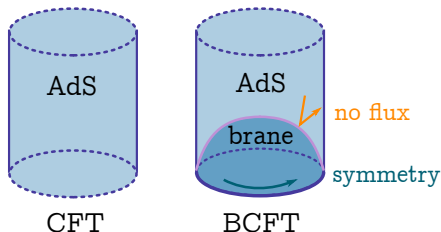
- RT formula states that **entanglement between A and complement \bar{A} is $\text{Area}(\gamma)/4G$** . Similar to black hole entropy!

[†]Van Raamsdonk 2009; Swingle 2009; Maldacena and Susskind 2013. [‡]Ryu and Takayanagi 2006; Hubeny, Rangamani, and Takayanagi 2007.

III. Boundary states

AdS/BCFT

- If we cut CFT in half, get a boundary CFT (BCFT).[†] Each half has **symmetry group $SO(d, 1)$ and leaks no energy.**
- Gravity dual is a **brane with same symmetry group and no flux (Neumann) boundary conditions.**[‡]

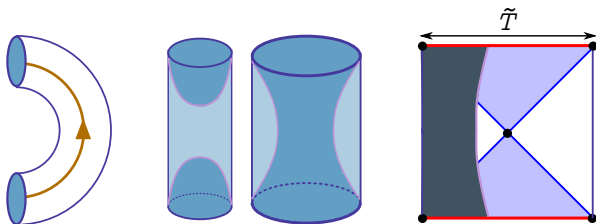


- $SO(d, 1)$ -symmetric BCFT configurations are called boundary (B) states $|B\rangle$.[§] **Model different $|B\rangle$ with brane tension \tilde{T} .**

[†]Cardy 1984. [‡]Karch and Randall 2001; Takayanagi 2011. [§]Cardy 1989.

Boundary state black holes

- Now consider BCFT at finite temperature. We **cut the donut in half** and get finite cylinder with two boundary components.
- **Two brane topologies**: disconnected and connected.

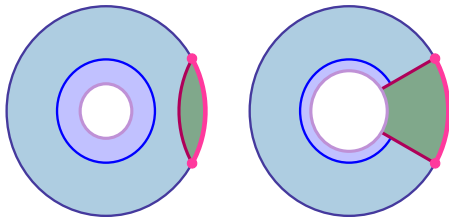


- **Connected phase is boundary state BH** (wormhole with brane).[†] It hits singularity at position determined by \tilde{T} .
- Get **state-dependent amount of interior**, as advertised!

[†]Fujita, Takayanagi and Tonni 2011; Almheiri, Mousatov, and Shyani 2018.

Hawking radiation and subsystem entanglement

- Can we decode Hawking radiation for boundary state BHs?
Too hard! **Entanglement of CFT regions is good surrogate.**
- Use RT formula. Two options for minimal surface: **outside horizon or onto brane[†]** when the brane is close enough.

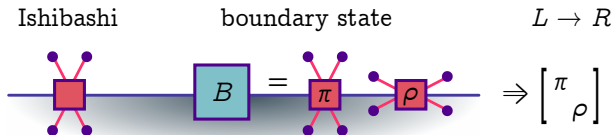


- Since brane size depends on tension \tilde{T} and time t , we can **decode boundary state $|B\rangle$ from subsystem entanglement.**

[†]Harlow 2016.

Boundary state simulation?

- Can we use **entanglement in B state to simulate interior?**
- Ishibashi states[†] (entangling left and right movers on either side of boundary) **satisfy no flux condition.**

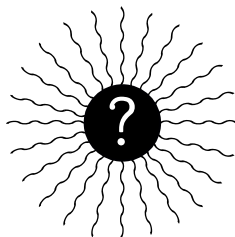


- **Combining Ishibashi states to preserve symmetry gives B states. B state is twisted map between left and right sectors!**
- Work in progress, but **twisted map can construct interior.**[‡]

[†]Ishibashi 1989. [‡] Almheiri 2018.

Loose threads

- Can we do **braneworld cosmology**?[†] Perhaps in charged BH!
- **Enlarge AdS/BCFT dictionary** to understand brane dynamics.
- Compare to **rigorous entropy calculations in BCFT₂**.[‡]
- Finally, see if B states **give insights into BHs or AdS/CFT**.[§]



Thanks for listening! Questions?

[†]Randall and Sundrum 1999; Karch and Randall 2000; Hebecker and March-Russell 2001. [‡]Cardy and Calabrese 2009. [§]Almheiri 2018.